

Feedbacks from user-oriented tools: CASSIS

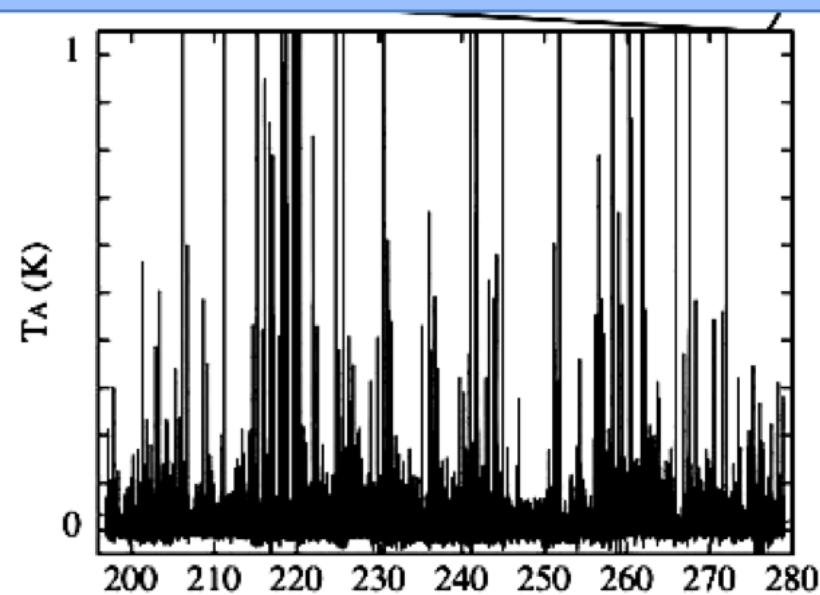
Charlotte Vastel

for the CASSIS team:

Jean-Michel Glorian & Mickael Boiziot

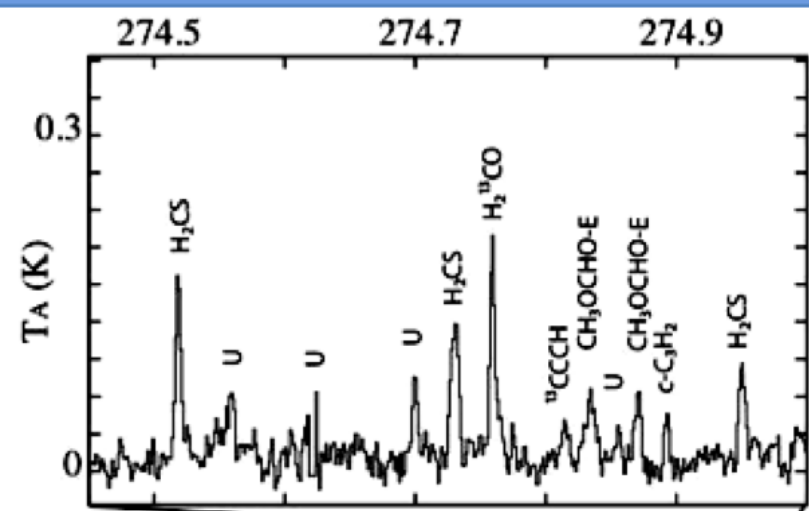
Emmanuel Caux & Sandrine Bottinelli

OBSERVATIONS



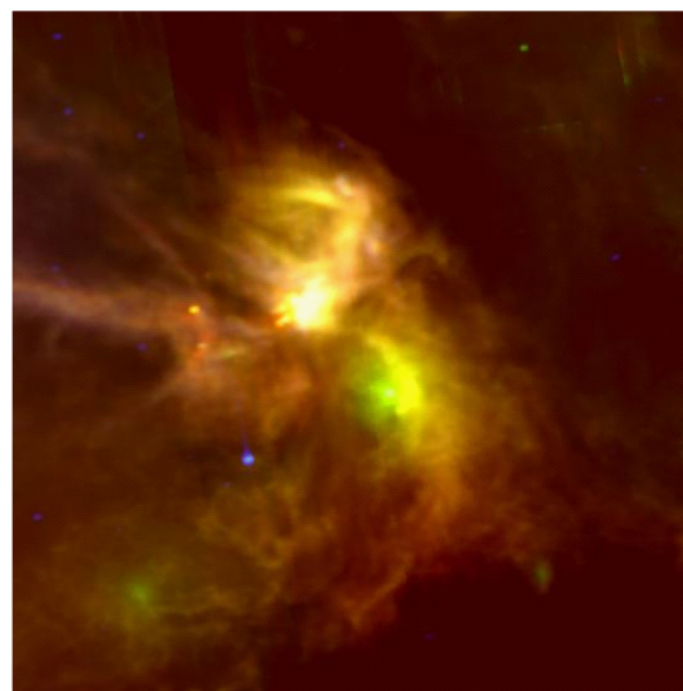
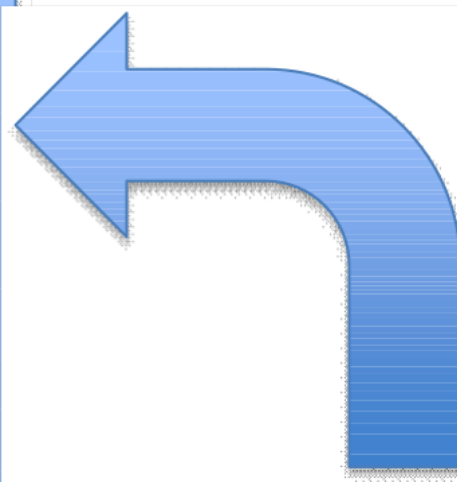
STEP1: Observe the spectrum of the source.

Tool: telescope

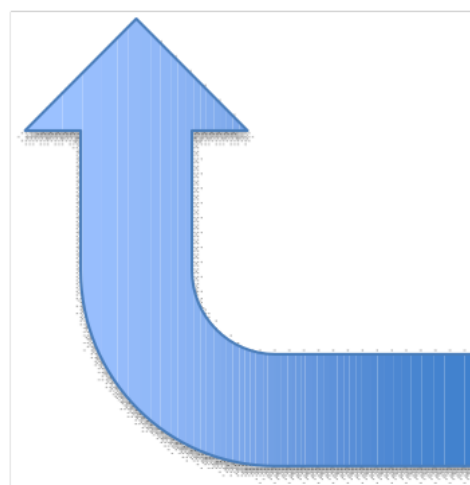


STEP2: Identify the lines and species.

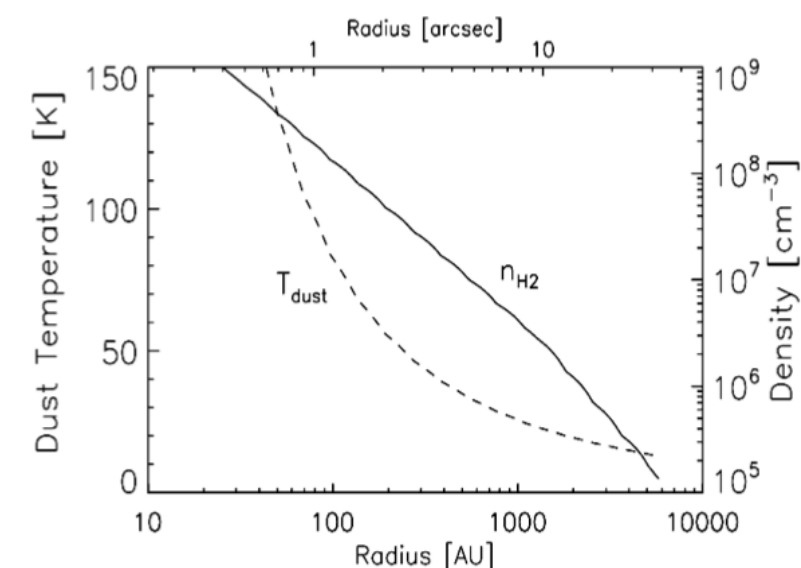
Tool: spectroscopic data



ASTROPHYSICAL
OBJECT

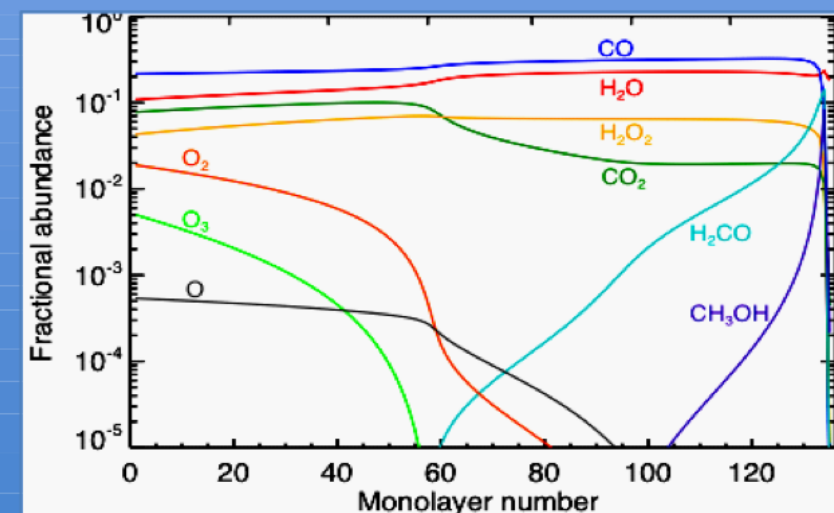


TOOLS & MODELS



STEP3: Derive the physical and chemical structure.

Tool: radiative transfer models

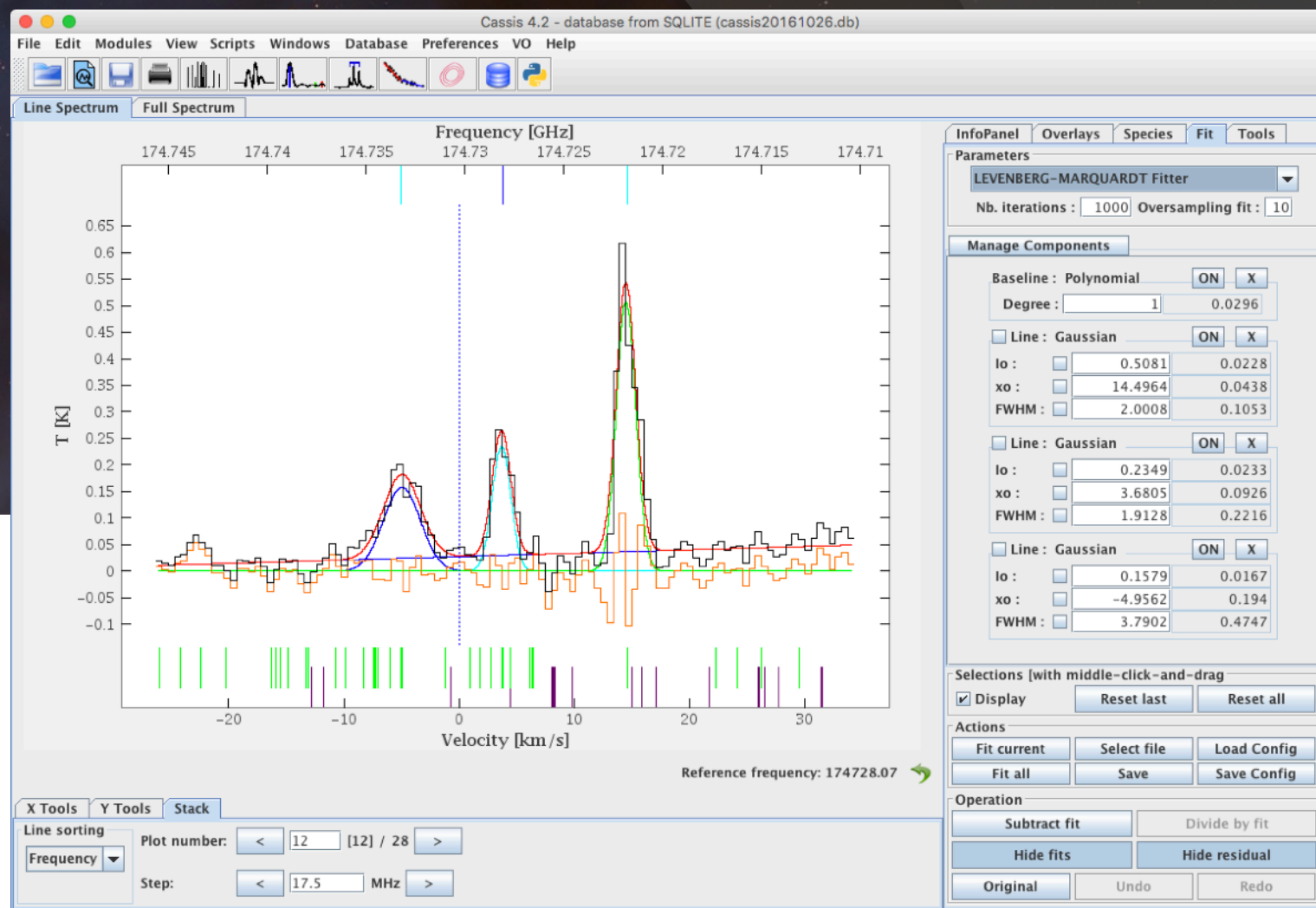


STEP4: Understand the chemical structure.

Tool: astrochemical models

CASSIS (Centre d'Analyse Scientifique de Spectres Instrumentaux et Synthétiques):

<http://cassis.irap.omp.eu/>



Astrophysical template

(fixed parameters N , T_k , T_{ex} , N_{H_2} , Δv , choice of the molecule...)

Observed spectra

(laboratory or telescope)
SSAP protocol (IVOA)

CASSIS

→ LTE model and Radex

→ Parameters to vary: N , T_k , T_{ex} , n_{H_2} , Δv , choice of the molecule and telescope, beam dilution...

Spectroscopic and molecular databases (JPL, CDMS, NIST)

TAP protocol (IVOA)

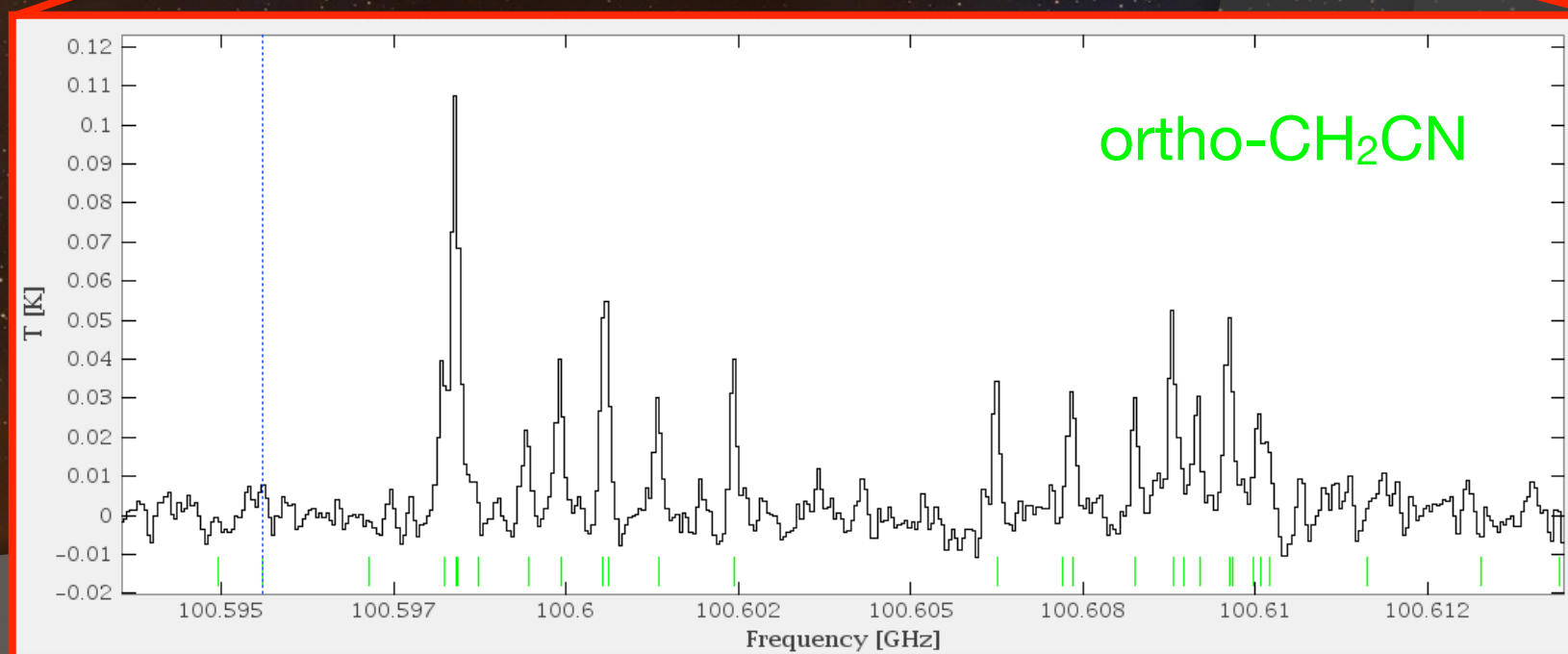
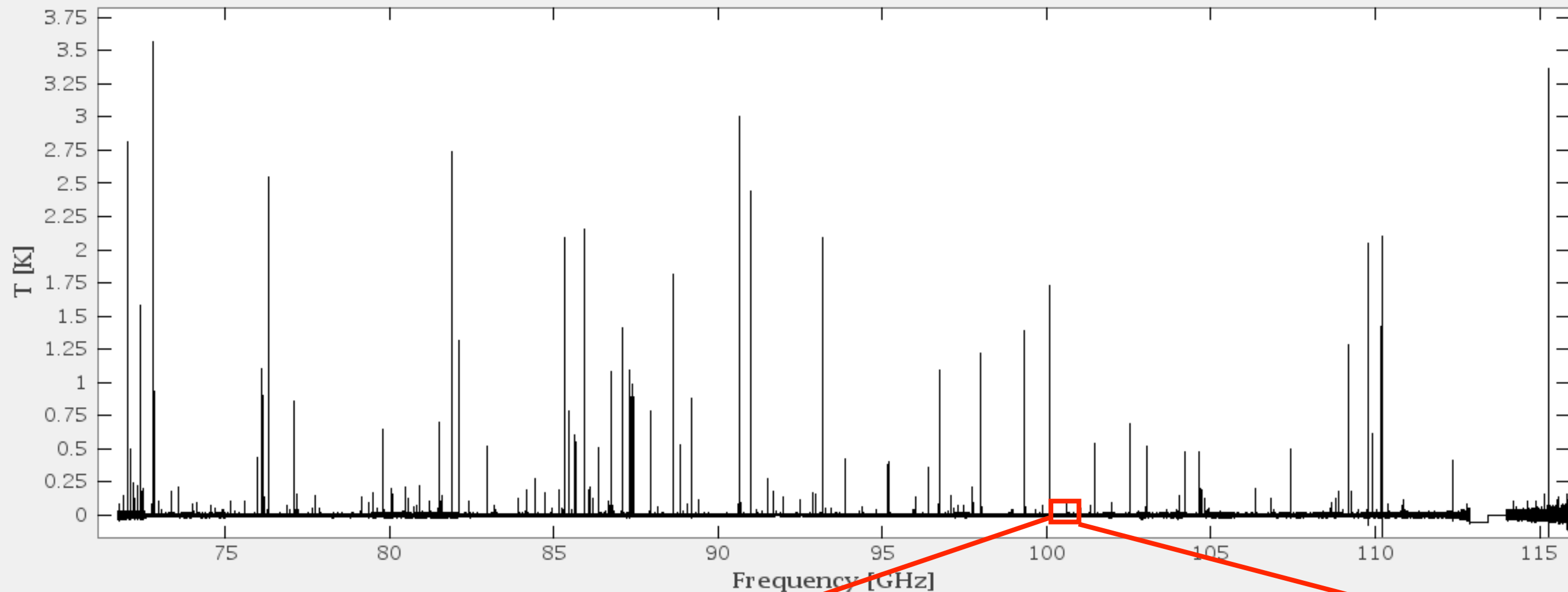


Synthetic spectra, Line identification, Automatic adjustment (MCMC, χ^2), Rotational diagram

+ CASSIS own databases with:

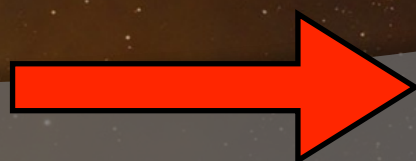
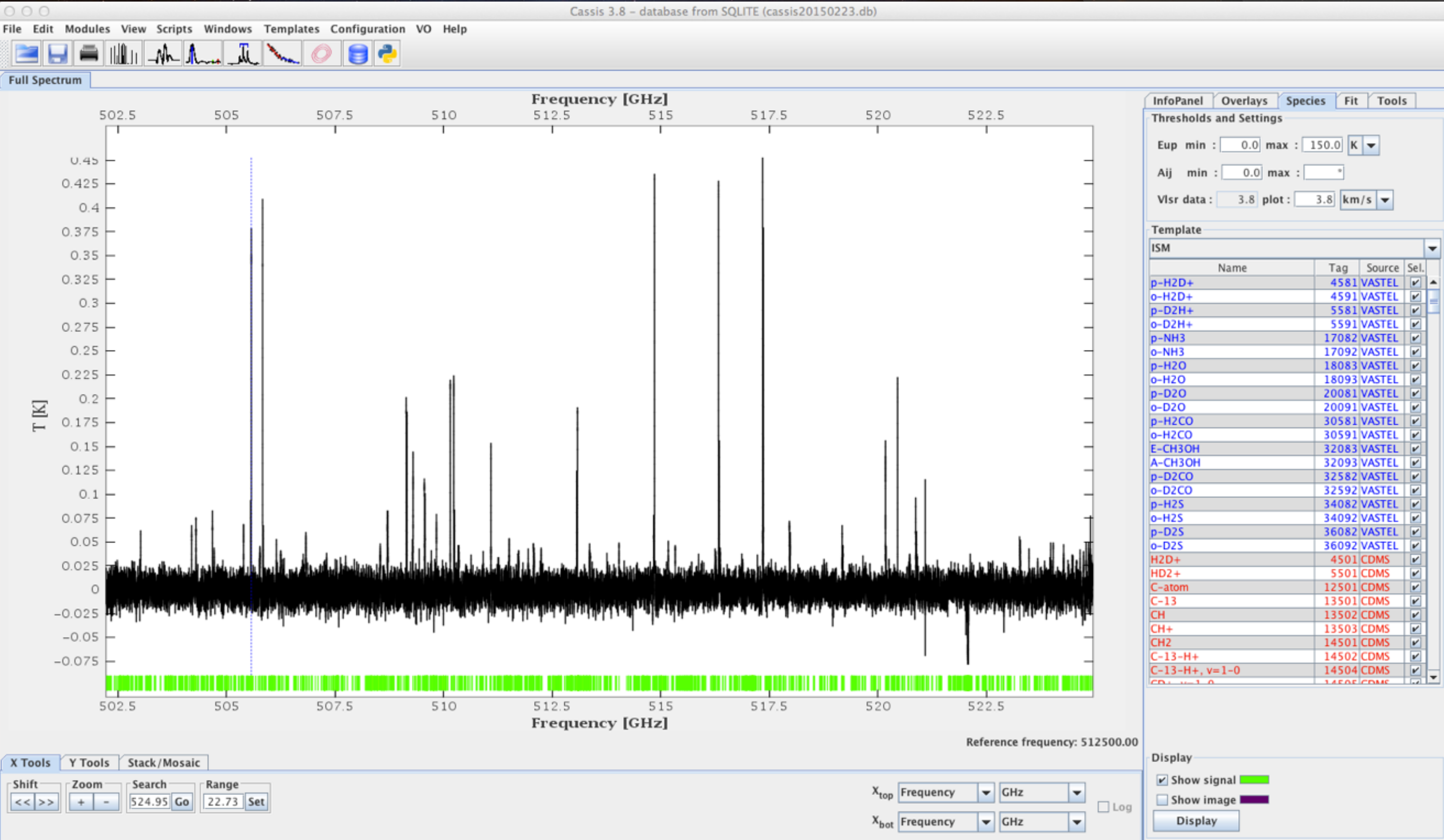
- 1) ortho / para / A / E separation
- 2) HFS
- 3) ~80 entries for Radex compatible collisional files

Line Identification



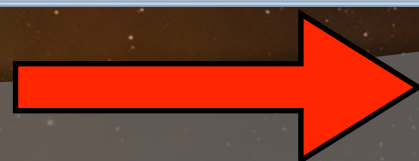
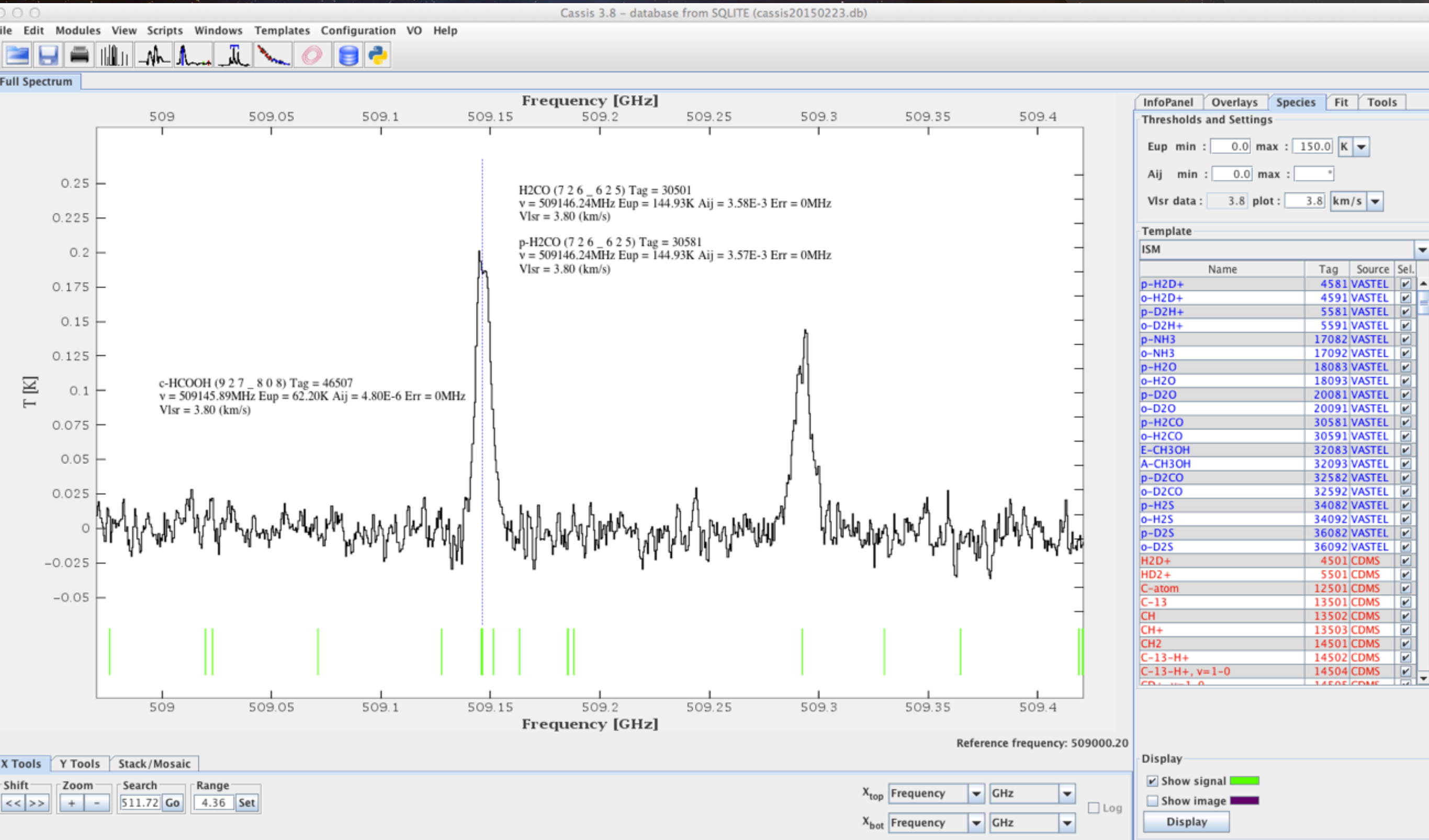
- * JPL
- * CDMS
- * NIST/VALD
- * Your own database
- * VAMDC

Line Identification



SQLite database

Line Identification



SQLite database

Line Identification

Databases and interoperability

The astronomers need: atomic, molecular databases



Local database (SQLite), built on CDMS, JPL, NIST, and private databases (lab or computations), database with nuclear spin state (ortho, para, A and E): very quick
CASSIS sometimes finds incoherences in the databases

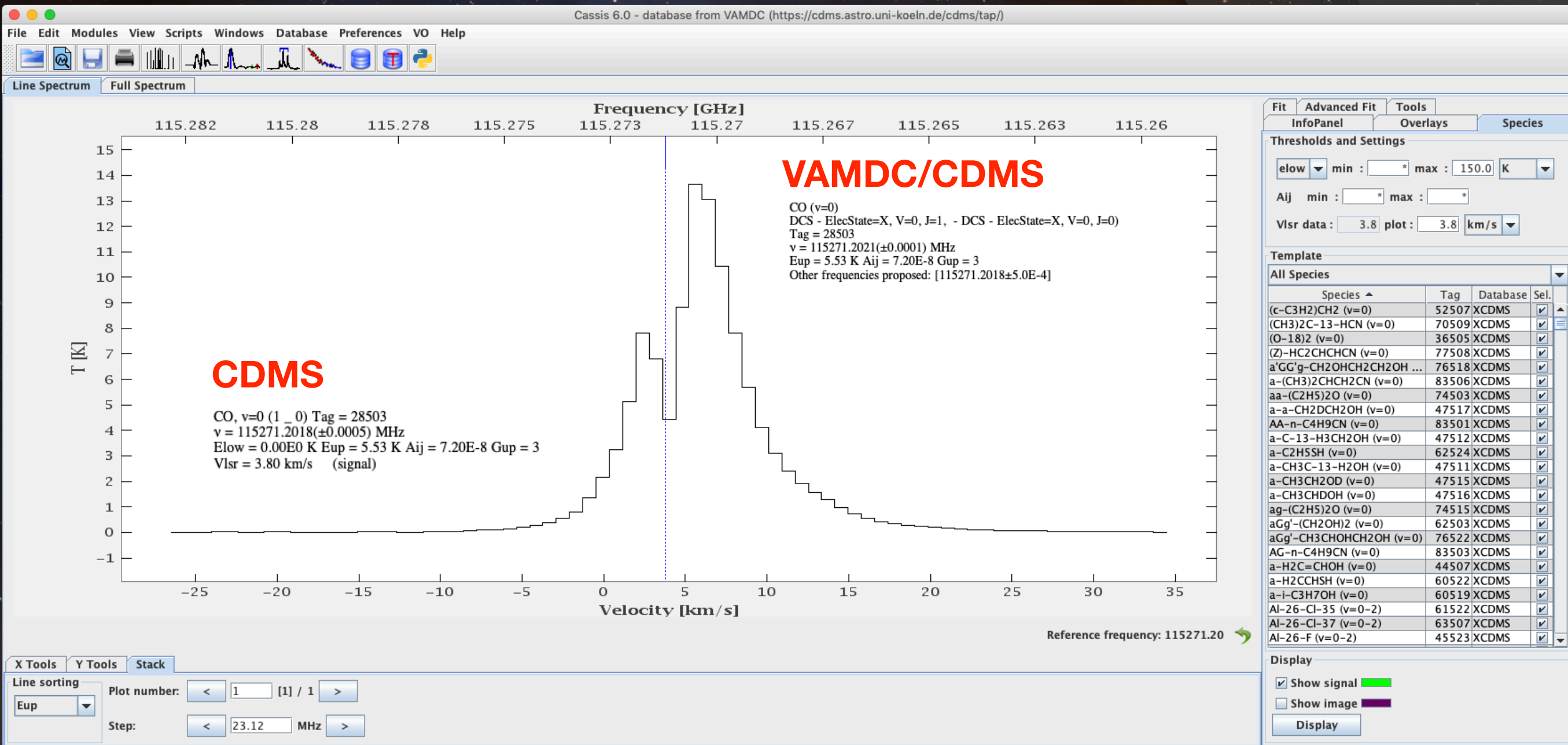


Access to the CDMS, JPL and VALD databases linked through VAMDC: very slow

Questions:

- 1) How is the update done through VAMDC? Regularly?
- 2) Within CASSIS, we must retrieve everything, references included: not useful! Too time consuming. Solution: json?
- 3) Through the portal, is it possible to select a database only?
And define criteria?

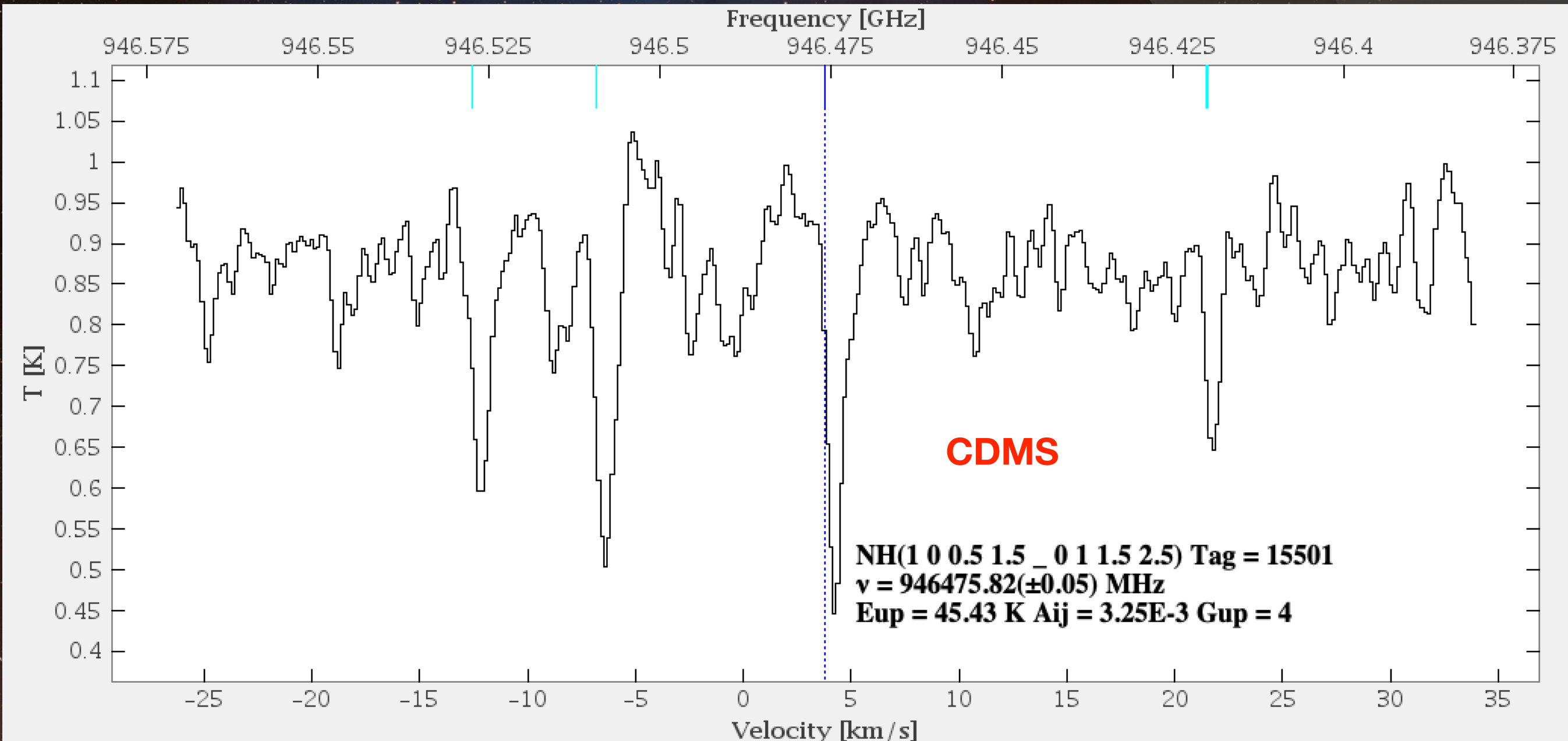
Line Identification: VAMDC Access to CDMS



When many frequencies available, we choose the one one with the lowest uncertainty.

VAMDC: not limited for example for the 3 digit limit on the statistical weigh

Line Identification: VAMDC Access to CDMS



NH from VAMDC decoded in CASSIS:

NH ($\nu=0$)

HUNDB - ElecState=X, S=1.0, F1NuclearSpinRef=N, V=0, FValue=1.5, J=0.0, F1Value=0.5, FNuclearSpinRef=H, N=1, -

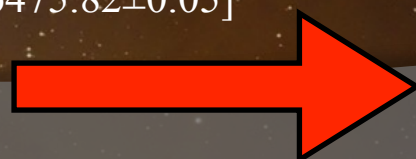
HUNDB - ElecState=X, S=1.0, F1NuclearSpinRef=N, V=0, FValue=2.5, J=1.0, F1Value=1.5, FNuclearSpinRef=H, N=0,)

Tag = 15501

$\nu = 946475.8538(\pm 0.033)$ MHz

Eup = 45.43 K Aij = 3.25E-3 Gup = 4

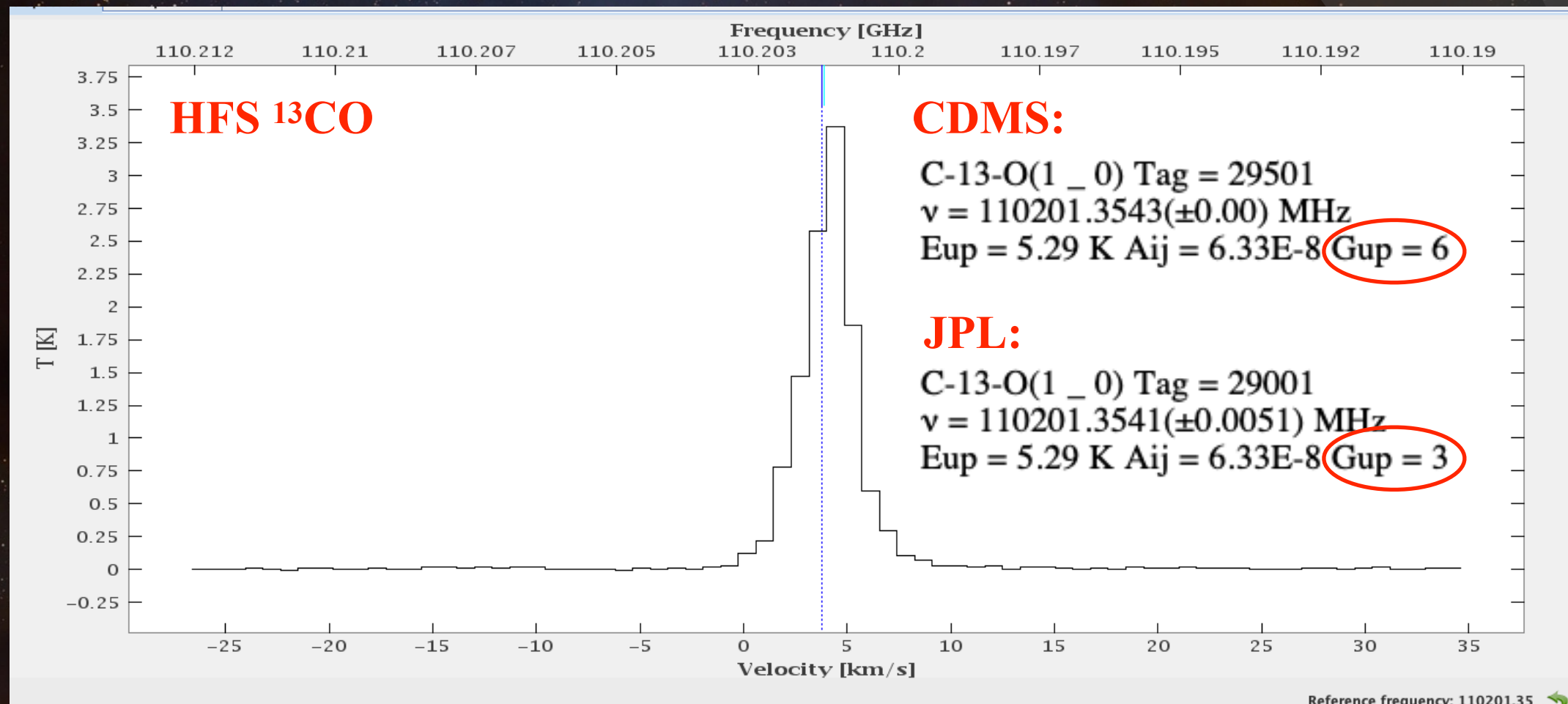
Other frequencies proposed: [946475.82 $\pm$ 0.05]



4 quantum numbers in CDMS

NH in VAMDC (15mn to download in VAMDC portal):

Line Identification: VAMDC Access to CDMS



VAMDC/CDMS

● DCS - ElecState=X, V=0, FValue=1.5, J=1, FNuclearSpinRef=C-13, - DCS - ElecState=X, V=0, FValue=0.5, J=0, FNuclearSpinRef=C-13
 $\nu = 110201.3707(\pm 0.0001) \text{ MHz}$, Eup = 5.29 K Aij = $6.33\text{E-}8$ Gup = 4

Other frequencies proposed: $[110201.37039 \pm 6.6\text{E-}4]$

● DCS - ElecState=X, V=0, J=1, - DCS - ElecState=X, V=0, J=0,
 $\nu = 110201.3543(\pm 0.00) \text{ MHz}$, Eup = 5.29 K Aij = $6.33\text{E-}8$ Gup = 6

● DCS - ElecState=X, V=0, FValue=0.5, J=1, FNuclearSpinRef=C-13, - DCS - ElecState=X, V=0, FValue=0.5, J=0, FNuclearSpinRef=C-13
 $\nu = 110201.3216(\pm 0.0002) \text{ MHz}$, Eup = 5.29 K Aij = $6.33\text{E-}8$ Gup = 2

Other frequencies proposed: $[110201.3218 \pm 7.0\text{E-}4]$

VAMDC/JPL:

● DCS - ElecState=X, V=0, J=1, - DCS - ElecState=X, V=0, J=0,)

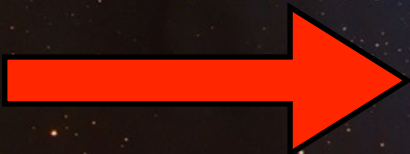
$\nu = 110201.3541(\pm 0.0051) \text{ MHz}$

Eup = 5.29 K Aij = $6.33\text{E-}8$ Gup = 3



Radiative transfert should be easier through VAMDC, but however much slower

Line Identification: Atomic databases

 **1 simple species, many different frequencies and A_{ij} , f_{ji} , S_{ik}**

CHIANTI: 1 4 1215.670 5.550e-01 6.260e+08 1s 2S1/2 - 2p 2P3/2 Lyman alpha

Chianti/VAMDC:

```
<SpeciesRef>Xchianti-1</SpeciesRef><Probability><TransitionProbabilityA><Value units="1/s">626000000.0</Value></TransitionProbabilityA><WeightedOscillatorStrength><Value units="unitless">0.555</Value></WeightedOscillatorStrength></Probability>
```

```
<ProcessClass></ProcessClass></RadiativeTransition><RadiativeTransition id="Pchianti-R10"><EnergyWavelength><Wavelength methodRef="Mchianti-EXP"><Value units="A">1215.67</Value></Wavelength><Wavelength methodRef="Mchianti-THEO"><Value units="A">1215.02</Value></Wavelength><EnergyWavelength><UpperStateRef>Schianti-3000001</UpperStateRef><LowerStateRef>Schianti-1000001</LowerStateRef>
```

NIST:

Observed	Ritz Wavelength	A_{ij}												
1215.6699	1215.668237310	6.2648e+08	AAA	0.0000000000-2	259.2850014	1s	2S	1/2	2p	2P°	3/2			

NIST/VAMDC: ATTENTION: There was an error in making the XML output and at least one item in the following parts was skipped: Atoms RadTran

VALD/VAMDC:

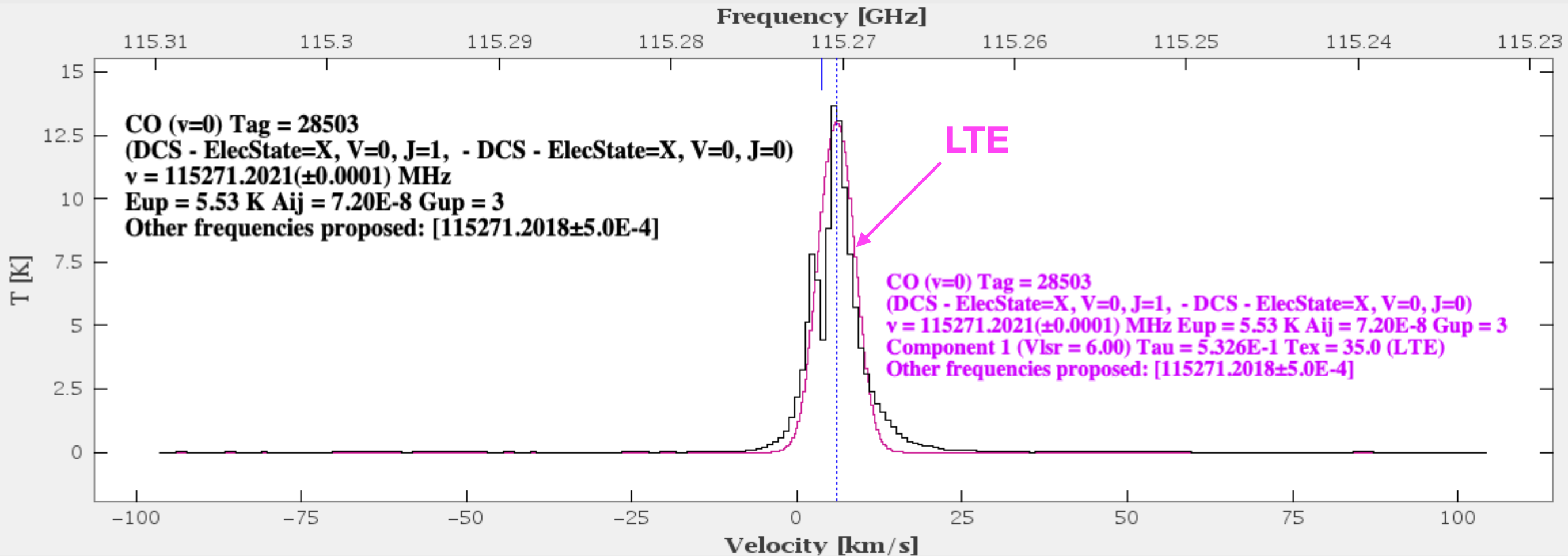
```
<SpeciesRef>Xvald-1</SpeciesRef><Probability><Log10WeightedOscillatorStrength><SourceRef>Bvald-CDROM18</SourceRef><Value units="unitless">-0.801</Value></Log10WeightedOscillatorStrength></Probability><ProcessClass></ProcessClass><Broadening name="natural" envRef="Evald-natural"><Comments>Natural Broadening</Comments><SourceRef>Bvald-CDROM18</SourceRef><Lineshape name="lorentzian"><LineshapeParameter name="log(gamma)"><Value units="1/s">8.77</Value></LineshapeParameter></Lineshape></Broadening></RadiativeTransition><RadiativeTransition id="Pvald-R154476610" process="excitation"><EnergyWavelength><Wavelength><Comments>Vacuum wavelength from state energies (RITZ)</Comments><SourceRef>Bvald-CDROM18</SourceRef><Value units="A">1215.67100000</Value><Wavelength><Wavelength><Comments>Vacuum wavelength from measurements (non-RITZ)</Comments><SourceRef>Bvald-CDROM18</SourceRef><Value units="A">1215.67089746</Value></Wavelength></EnergyWavelength><UpperStateRef>Svald-1105790</UpperStateRef><LowerStateRef>Svald-1021407</LowerStateRef>
```

Kentucky database (Not in VAMDC):

-LAMBDA-VAC-ANG-	-SPECTRUM--	TT	-----TERM-----	---J-J---	----LEVEL ENERGY--CM-1----
1215.6700	H I	E1	1-2	1/2-*	0.00 - 82259.11

Radiative transfer: LTE

- 1) Line identification through VAMDC
- 2) LTE modelling using the partition function, A_{ij} , g_u , frequencies given by VAMDC



Radiative transfer: non-LTE

Collisional databases

Line Analysis 1

Data
Load IRAS16293_SS/iram.bas Vlsr data: 3.8 km/s in: REST Telescope ???

Tuning
Range min: 79.9948749 max: 281.004874 GHz Band: 60.0 km/s

Threshold
Eup min: 0.0 max: 150.0 K Aij min: 0.0 max: *
Jup min: * max: * Kup min: * max: * Lup min: * max: * Mup min: * max: *

LTE-RADEX ☒

Telescope
apex ☐ Tmb->Ta *

Observing Mode
PSw/DBSw

Background
Tbg: 2.73 K

Noise
rms: 0.0 mK

Oversampling
Oversampling: 3.0

Continuum
☒ Component 1 X

Mode: Full Radex
Molecules: -- Operations --

☒ Interacting
Geometry: Sphere

$N(H_2)$ [cm⁻²]: 7.5E22
 V_{lsr} : 3.8 km/s

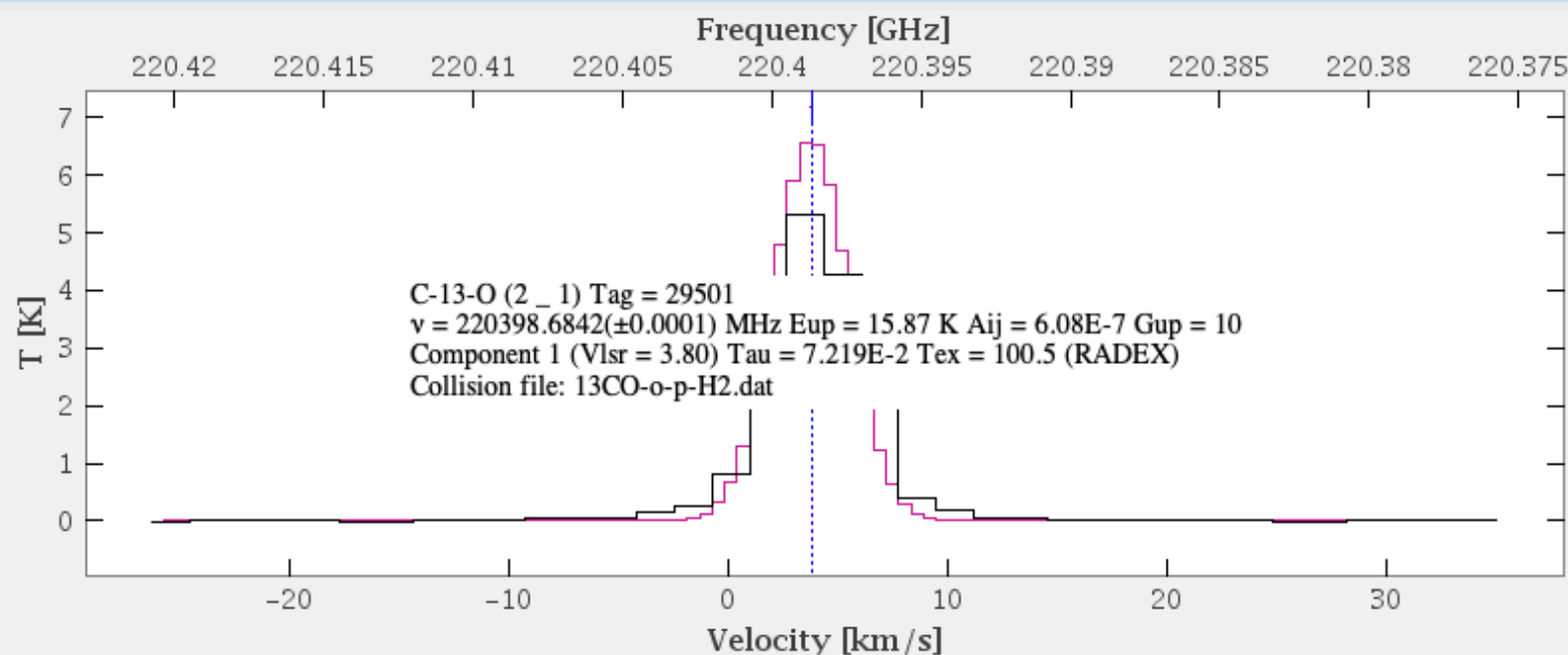
Species	Tag	Database	Collision	Compute	N(Sp) (cm ⁻²)	Abundance (/...)	TKin (K)	FWHM (km/s)	Size (")
C-13-O	29501	CDMS	13CO-o-p-H...	☒	4.00E16	5.33E-7	100.00	4.00	3000.00

Cassini 6.0 - database from SQLITE (cassini20200513-FORCOM-HFS-IRAP.db)

File Edit Modules View Scripts Windows Database Preferences VO Help



Line Spectrum



Fit Advanced Fit Tools
InfoPanel Overlays Species

Thresholds and Settings

eup min : * max : 150.0 K
Aij min : * max : *
Vlsr data : 3.8 plot : 3.8 km/s

Template

All Species			
Species	Tag	Database	Sel.
(c-C3H2)CH2	52507	CDMS	☐
(O-18)2	36505	CDMS	☐
13C-CCH, v4=0,1mS	38504	CDMS	☐
13CH	14003	JPL	☐
13CH313CN	43005	JPL	☐
13CH3CH2OH	47094	IRAP	☐
13CH3D	18006	JPL	☐
15-NO	31009	JPL	☐

Display

☐ Show signal
☐ Show image

Display

X Tools Y Tools Stack

Line sorting

Frequency

Plot number: < 2 [1,2] / 2 >

Step: < 22.19 MHz >

!MOLECULE

HCN

!MOLECULAR WEIGHT

27.0

!NUMBER OF ENERGY LEVELS

4

!LEVEL + ENERGIES(cm⁻¹) + WEIGHT + J,F

1	0.0000	3.0	0 1
2	2.9564	1.0	1 0
3	2.9564	3.0	1 1
4	2.9564	5.0	1 2

!NUMBER OF RADIATIVE TRANSITIONS

33

!TRANS + UP + LOW + EINSTEINA(s⁻¹) + FREQ(GHz) + E_u(K)

1	2	1	2.4075e-05	88.6339360	4.25
2	3	1	2.4075e-05	88.6304160	4.25
3	4	1	2.4075e-05	88.6318470	4.25

!NUMBER OF COLL PARTNERS

1

!COLLISIONS BETWEEN

1 HCN-H₂ scaled (*1.37) HCN-He from Green & Thaddeus (1974) + extrapolation

!NUMBER OF COLL TRANS

6

!NUMBER OF COLL TEMPS

4

!COLL TEMPS

			5.0	10.0	20.0	30.0
!TRANS + UP + LOW + COLLRATES(cm ³ s ⁻¹)						
1	2	1	1.3e-11	1.1e-11	9.2e-12	8.4e-12
2	3	1	1.3e-11	1.1e-11	9.2e-12	8.4e-12
3	3	2	0.0e+00	0.0e+00	0.0e+00	0.0e+00
4	4	1	1.3e-11	1.1e-11	9.2e-12	8.4e-12
5	4	2	0.0e+00	0.0e+00	0.0e+00	0.0e+00
6	4	3	0.0e+00	0.0e+00	0.0e+00	0.0e+00

**Spectroscopic parameters:
matching needs to be done
between the collision files
and the spectroscopic
parameters (JPL, CDMS)
used in CASSIS.**

**For 1 molecule found in
CDMS and JPL we need 2
collisional files (different
quantum numbers,
frequencies etc...)**

Collisions


```
! Molecule
H2O - ortho
! MASS
18.01
! Number of Energy Level : jpl V6 PNov. 2005 -- http://spec.jpl.nasa.gov/ftp/pub/catalog/doc/d018003.cat
5
! LEVEL + ENERGY(CM-1) + WEIGHT + QUANTUM NOS. Ka_Kc_tau_J
 1  23.794400    9  0_1_-1_1  —————→ 4
 2  42.371700    9  1_0_1_2
 3  79.496400   15  1_2_-1_2
 4  134.901600   15  2_1_1_3
 5  136.761700   21  0_3_-3  —————→ 3
! NUMBER OF RADIATIVE TRANSITIONS : jpl V6 PNov. 2005 -- http://spec.jpl.nasa.gov/ftp/pub/catalog/doc/d018003.cat
10
! TRANS + UP + LOW + EINSTEINA(s^-1) + FREO(GHz) + E_up(K)
 1  2  1  34.6009E-4  556.9360020  60.966475  —————→ ?
 2  2  1  34.6009E-4  556.9360020  60.966576
 3  3  1  5.59938E-2  1669.9047750  114.383309
 4  3  1  5.59938E-2  1669.9047750  114.383273
 5  4  2  25.7413E-2  2773.9765880  194.103022
 6  4  2  25.7413E-2  2773.9765880  194.103127
 7  4  3  3.06509E-2  1661.0076370  194.103127
 8  4  3  3.06509E-2  1661.0076370  194.103022
 9  5  3  5.05947E-2  1716.7696330  196.779424
10  5  3  5.05947E-2  1716.7696330  196.779452
! NUMBER OF COLLISION PARTNERS
1
! COLLISION PARTNER
1
! COLLISIONS BETWEEN, BASECOL V1 2017-10
2 para-H2-H2O-ortho
! NUMBER OF COLLISIONAL TRANSITIONS
10
! NUMBER OF COLLISION TEMPERATURES
7
! COLLISION TEMPERATURES
20.0  40.0  60.0  80.0  100.0  120.0  140.0
! TRANS + UP + LOW + RATE COEFFS(cm^3 s^-1)
 1  2  1  1.1700e-11  1.6000e-11  1.8900e-11  2.1600e-11  2.4400e-11  2.7200e-11  3.0300e-11
 2  3  1  2.0200e-11  2.1700e-11  2.2500e-11  2.3500e-11  2.4700e-11  2.6300e-11  2.7800e-11
 3  3  2  1.6800e-11  1.6800e-11  1.6100e-11  1.5600e-11  1.5200e-11  1.4900e-11  1.4800e-11
```

different quantum numbers compared to JPL

different frequencies and Einstein coef compared to JPL

References to the collisional database missing

JPL: 556935.9877 0.0003 -0.8189 3 23.7944 9 -180031404 1 1 0 0 1 0 1 0

Databases and interoperability

Our needs for radiative transfer: collisional databases based on LAMDA and/or BASECOL

The CASSIS software provides a database for the collision rates with a matching for the provided databases (~ 80 entries)

➔ One species, multiple IDs. Exemple: SO 48501 (CDMS) and 48001 (JPL). Multiple collision files (ortho and para H₂, He scaled, multiple authors).

➔ Matching of the frequencies, A_{ij} , quantum numbers between the collision files and the spectroscopic databases to produce synthetic spectra.

**To do for CASSIS: directly interrogate VAMDC
to use the collision files (XSAMS): SAMP?**



<http://cassis.irap.omp.eu/?page=catalogs-collision>

45506	HCS+	hcsp-h2.dat	H2	adopting HCO+ - H2 from Flower (1999) + extrapolation	from LAMDA, Aug 24, 2017	T=[10-2000] K
48001	SO	so-He-scaled.dat	He scaled (*1.363)	Lique et al (2006)	from LAMDA, Aug 23, 2017	T=[60-300] K
48001	SO	so-pH2.dat	para H2	Lique et al (2007)	constructed by CV from BASECOL, Aug 23, 2017	T=[5-50] K
48501	SO	so-He-scaled.dat	He scaled (*1.363)	Lique et al (2006)	from LAMDA, Aug 23, 2017	T=[60-300] K
48501	SO	so-pH2.dat	para H2	Lique et al (2007)	constructed by CV from BASECOL, Aug 23, 2017	T=[5-50] K
51501	HC3N	hc3n-oph2-cdms.dat	ortho and para H2	Faure et al. 2016 MNRAS 460 2103	Given by A. Faure	T=[10-300] K